

3.3. Renormalisation Group (RG)

(59)

- in order to regularise the 2-point function we have introduced a mass scale " μ ". In general the renormalisation factors Z_i will depend on μ , as we have seen for Z_m (and will see for Z_A below). Also m_R, λ_R will in general on μ (unlike m_B, λ_B).

But $G_B^{(n)}$ is indep of $\mu \Rightarrow$ the (dimensionless) derivative $\mu \frac{d}{d\mu} G_B^{(n)}(x_1, \dots, x_n; m_B, \lambda_B) = 0$ where we made the parameter-dependence explicit

From $G_B^{(n)}(x_1, \dots, x_n; m_B, \lambda_B) = Z_\phi^{\frac{n}{2}} G_R^{(n)}(x_1, \dots, x_n; m_R, \lambda_R)$ it follows

$$0 = \mu \frac{d}{d\mu} \left(Z_\phi^{\frac{n}{2}} G_R^{(n)} \right) = \frac{n}{2} \mu \frac{\partial Z_\phi}{\partial \mu} Z_\phi^{\frac{n}{2}-1} G_R^{(n)} + Z_\phi^{\frac{n}{2}} \left(\mu \frac{\partial}{\partial \mu} G_R^{(n)} + \mu \frac{\partial m_R}{\partial \mu} \frac{\partial}{\partial m_R} G_R^{(n)} + \mu \frac{\partial \lambda_R}{\partial \mu} \frac{\partial}{\partial \lambda_R} G_R^{(n)} \right)$$

$$\Leftrightarrow \left[\mu \frac{\partial}{\partial \mu} + \beta(\lambda_R) \frac{\partial}{\partial \lambda_R} + \gamma_m(\lambda_R) \frac{\partial}{\partial m_R} + n \gamma(\lambda_R) \right] G_R^{(n)} = 0$$

RG eq. for MS^t

with $\beta(\lambda_R) = \mu \frac{\partial \lambda_R}{\partial \mu}$, $\gamma_m(\lambda_R) = \mu \frac{\partial \ln m_R}{\partial \mu}$, $\gamma(\lambda_R) = \mu \frac{\partial \ln Z_\phi}{\partial \mu}$

The study of the β -function $\beta(\lambda_R)$ will be important later.

- these eqs can be formulated in Fourier space and/or on $\Gamma^{(n)}$ instead.
- the RG eqs. can be formulated differently:

Callan-Symanzik: $m_R \frac{\partial}{\partial m_R} G = \dots$ involving only R-quantities

Wilson RG: $\Lambda \frac{\partial}{\partial \Lambda} G = \dots$ in terms of a cutoff Λ (here IR!)

(see also Franz Wegner)

- RG-studies used also in phase-transitions in stat. Phys!

Regularisation of λ and determination of Z_1

Therefore we will consider $G_{3,0}^{(4)}$ to order λ^2 : from page (37) or using the path integral formalism we have

(a) $G_{3,0} \circlearrowleft(\lambda_0^0)$: $\langle \phi_1 \phi_4 \rangle_0$ or \equiv only disconnected

$\circlearrowleft(\lambda_0^1)$: $\langle \phi_1 \phi_4 [\phi^4] \rangle_0 = \langle \phi_1 \phi_4 \rangle_0 \langle [\phi^4] \rangle_0$ or $\frac{0}{0} + X$
 \Rightarrow only $\begin{matrix} 1 & 2 \\ 2 & 1 \end{matrix}$ $\begin{matrix} 1 & 2 \\ 2 & 1 \end{matrix}$ combact with X ; \uparrow discon, \uparrow connected

we get for $\circlearrowleft(\lambda_0^2)$:

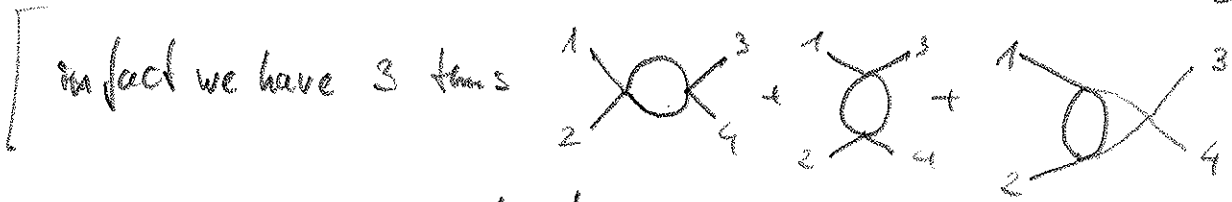
$\frac{1}{2} \langle \phi_1 \phi_4 [\phi^4] [\phi^4] \rangle_0 = 2 \langle \phi_1 \phi_4 [\phi^4] \rangle_0 \langle [\phi^4] \rangle_0$

$-\frac{1}{2} \langle \phi_1 \phi_4 \rangle_0 \langle [\phi^4] [\phi^4] \rangle_0 + \langle \phi_1 \phi_4 \rangle_0 \langle [\phi^4] \rangle_0^2$

\Rightarrow only two s connecty $\begin{matrix} 1 & 3 \\ 2 & 4 \end{matrix}$ with both X X



only connected



\Rightarrow Mandelstam variables for external momenta p_1, \dots, p_4

$s = (p_1 + p_2)^2, t = (p_2 + p_3)^2, u = (p_1 + p_4)^2$

• for the explicit computation we go to Fourier space

$(i\omega)^4 \delta(p_1 + p_2 + p_3 + p_4) G_{3,0}^{(4)}(p_1, p_2) = \frac{4!}{i^4} \int \prod_{i=1}^4 dx_i e^{i p_i x_i} G_{3,0}^{(4)}(x_1, \dots, x_4)$

and set the external momenta $p_1 = \dots = p_4 = 0$ (this is all we need)

• for $p_i \neq 0$ one can use the Feynman parametrisation \rightarrow ext

$$\begin{aligned}
 \underline{O(\lambda_B^1)}: \text{in } \delta(0) \left. G_{B,C}^{(4)}(0, \dots, 0) \right|_{\lambda_B} &= \int_{k_1, \dots, k_4} \int_w \frac{e^{i0 \cdot x}}{w} \left(\frac{-\lambda_B}{4!} \right) \langle \phi_B(k_1) \phi_B(k_2) \phi_B(k_3) \phi_B(k_4) \phi_B(w) \phi_B(w) \rangle_0 \\
 &= -\lambda_B \int_{k_1, \dots, k_4} \int_w \frac{e^{i p_1(k_1-w)}}{p_1^2 + m_B^2} \cdot \frac{e^{i p_2(k_2-w)}}{p_2^2 + m_B^2} \cdot \frac{e^{i p_3(k_3-w)}}{p_3^2 + m_B^2} \cdot \frac{e^{i p_4(k_4-w)}}{p_4^2 + m_B^2} \\
 &\quad \uparrow \quad \uparrow \\
 &\quad \frac{4}{\pi} \delta(p_i) \quad \delta(p_1 + p_2 + p_3 + p_4) \\
 &= -\lambda_B \delta(0) \left(\frac{1}{m_B^2} \right)^4 \quad \text{to } \mathcal{O}(\lambda_B) \quad \times \quad \sim \lambda_B
 \end{aligned}$$

$O(\lambda_B^2)$:

$$\text{in } \delta(0) \left. G_{B,C}^{(4)}(0, \dots, 0) \right|_{\lambda_B^2} = \int_{k_1, \dots, k_4} \int_{w, v} \frac{1}{2} \frac{\lambda_B^2}{(4!)^2} \langle \underbrace{\phi_B(k_1) \phi_B(k_2) \phi_B(k_3) \phi_B(k_4)}_{\text{etc}} \phi_B(w)^4 \phi_B(v)^4 \rangle_0$$

connected, only contr. with 4 w and v

$$= \dots = \frac{3}{2} \lambda_B^2 \delta(0) \left(\frac{1}{m_B^2} \right)^4 \int \frac{d^d p}{(2\pi)^d} \frac{1}{(p^2 + m_B^2)^2}$$

divergent, regularise

in $d = 4 - 2\epsilon$

$$\int \frac{d^d p}{(2\pi)^d} \frac{1}{p^2 + m^2} = -\frac{\mu^{-2\epsilon} m^2}{(4\pi)^2} \left(\frac{1}{\epsilon} + 1 - \gamma_E + \ln \frac{\mu^2}{m^2} + \ln(4\pi) + \mathcal{O}(\epsilon) \right)$$

\Rightarrow

$$\int \frac{d^d p}{(2\pi)^d} \frac{1}{(p^2 + m^2)^2} = -\frac{d}{dm^2} \int \frac{d^d p}{(2\pi)^d} \frac{1}{p^2 + m^2} = -\frac{\mu^{-2\epsilon}}{(4\pi)^2} \left(\frac{1}{\epsilon} + \ln \frac{\mu^2}{m^2} + \mathcal{O}(\epsilon) \right)$$

• using $\lambda_B = (1 + \delta Z_\lambda) \lambda_R$ (and the fact that $Z_\phi = 1 + \mathcal{O}(\lambda_R^2)$ so that it can be neglected here in $G_{B,C}^{(4)} = Z_\phi^2 G_{R,C}$)

we find

$$\begin{aligned}
 \left(m_B^2 \right)^4 \delta(0) \left. G_{R,C}^{(4)}(0, \dots, 0) \right|_{\lambda_R} &= \lambda_{\text{phys}} = \lambda_R + \lambda_R \delta Z_\lambda - \frac{3}{2} \frac{\lambda_R^2 \mu^{-2\epsilon}}{(4\pi)^2} \left(\frac{1}{\epsilon} + \ln \frac{\mu^2}{m_R^2} + \mathcal{O}(\epsilon) \right) \\
 \uparrow \\
 \text{removing ext. legs} & \quad + \mathcal{O}(\lambda_R^3)
 \end{aligned}$$

o in the term of order Λ_B^2 we can replace $B \rightarrow R$ everywhere to L.O.

\Rightarrow in the MS scheme we choose
$$\boxed{Z_\lambda = \frac{3}{2} \frac{\Lambda_R \mu^{-2\epsilon}}{(4\pi)^2} \frac{1}{\epsilon}}$$

Determination of $\beta(\lambda) = \mu \frac{\partial \Lambda_R}{\partial \mu}$: (in L.O. in pert. theory)

$$\Lambda_B = \Lambda_R Z_\lambda = \Lambda_R + \frac{3}{2} \frac{\Lambda_R^2 \mu^{-2\epsilon}}{(4\pi)^2} \frac{1}{\epsilon} + O(\Lambda_R^3)$$

$$\Rightarrow 0 = \mu \frac{d}{d\mu} \Lambda_B = \mu \frac{\partial \Lambda_R}{\partial \mu} + \frac{3}{2} \frac{\Lambda_R^2}{(4\pi)^2} \underbrace{\mu \frac{\partial}{\partial \mu} \frac{\mu^{-2\epsilon}}{\epsilon}}_{-2\mu^{-2\epsilon}} + \frac{3}{2} \frac{\mu^{-2\epsilon}}{(4\pi)^2} \underbrace{2\Lambda_R \mu \frac{\partial \Lambda_R}{\partial \mu}}_{\text{if we retract the eq this term is of } O(\Lambda_R^3)}$$

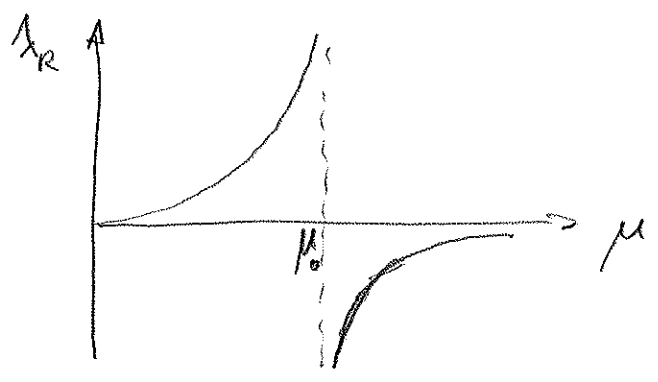
\Rightarrow to L.O.
$$\boxed{\mu \frac{\partial \Lambda_R}{\partial \mu} = \frac{3}{(4\pi)^2} \Lambda_R^2}$$

- the derivation of suchs RGE equations is simple once the Z_i - factors are known
- the eqs are finite in the limit $\epsilon \rightarrow 0$ as these are eqs. for renormalised quantities :

Solution (ex 7.4) for $\Lambda_R(\mu)$ to this order ..

$$\Lambda_R(\mu) = \frac{(4\pi)^2}{3} \frac{1}{\ln \frac{\mu_0}{\mu}}$$

with some (dim-ful) constant μ_0 .

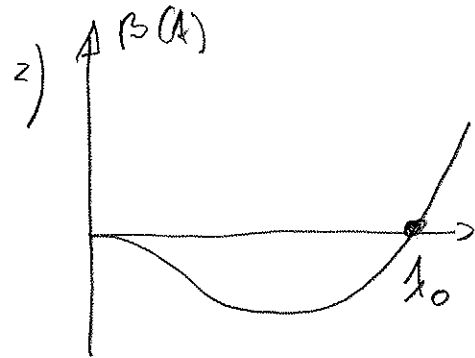
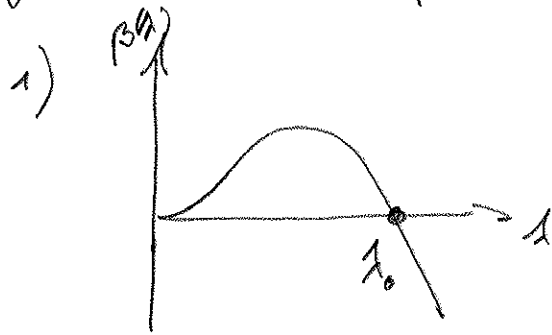


We can only trust this for $\Lambda_R \ll 1$ (and $\mu \ll \mu_0$)

Fixed points of the β -function

We had defined $\beta(\lambda_R) = \mu \frac{\partial \lambda_R}{\partial \mu}$. Let us discuss 2

generic scenarios for $\beta(\lambda)$:



(One could hope that higher orders remove the pole in $\lambda_R(\mu)$ and map ϕ -theory to scenario 1) - this seems only be true for $d=3$)

• λ_0 is a fixed point (FP) of $\beta(\lambda)$ - is it stable?

around $\lambda \approx \lambda_0$ we can linearise:

$$\begin{aligned} \beta(\lambda) &\approx -B(\lambda - \lambda_0) && \text{with } B > 0 \text{ in 1)} \\ &= \mu \frac{\partial \lambda}{\partial \mu} && B < 0 \text{ in 2)} \end{aligned}$$

\Rightarrow solution $\lambda(\mu) = \lambda_0 + e^{-B \ln \mu} = \lambda_0 + \frac{1}{\mu^B}$

\Rightarrow the scale dependent (running) coupling $\lambda(\mu)$ is driven

towards λ_0 for $B > 0$ $\lim_{\mu \rightarrow \infty} \left(\lambda(\mu) = \lambda_0 + \frac{1}{\mu^B} \right) = \lambda_0$ 1)

$B < 0$ $\lim_{\mu \rightarrow 0} \left(\lambda(\mu) = \lambda_0 + \mu^{|B|} \right) = \lambda_0$ 2)

Therefore in 1) we have a UVFP

2) we have an IR FP - this is possible in an asymptotically free theory